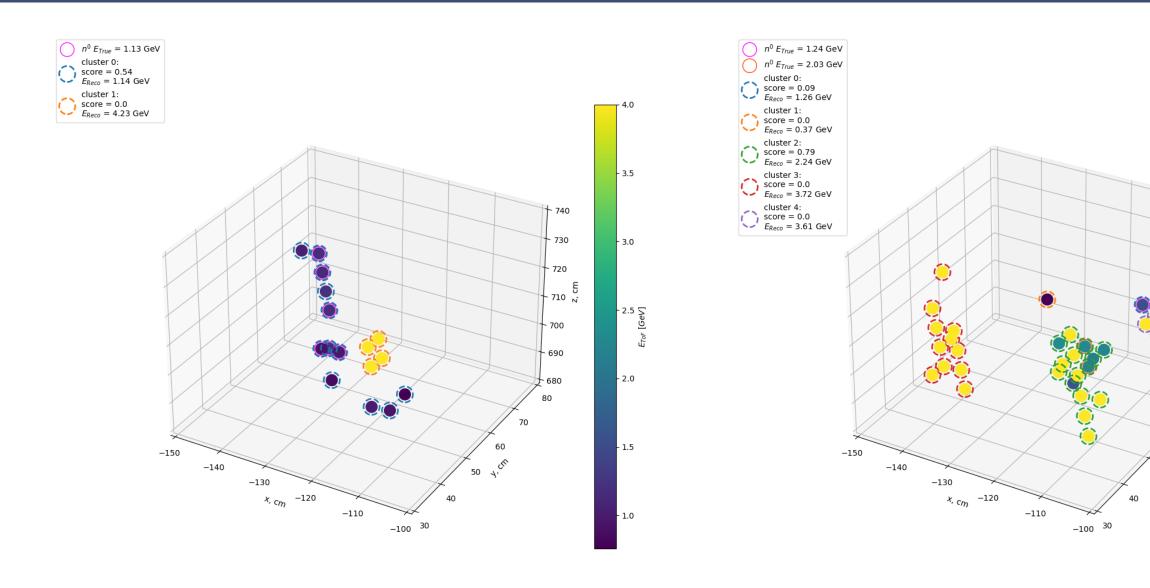
# GNN-based neutron reconstruction in the highly-granular neutron detector at the BM@N experiment

#### Vladimir Bocharnikov

29.10.2025

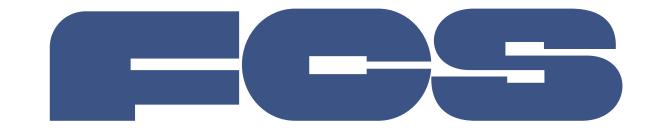










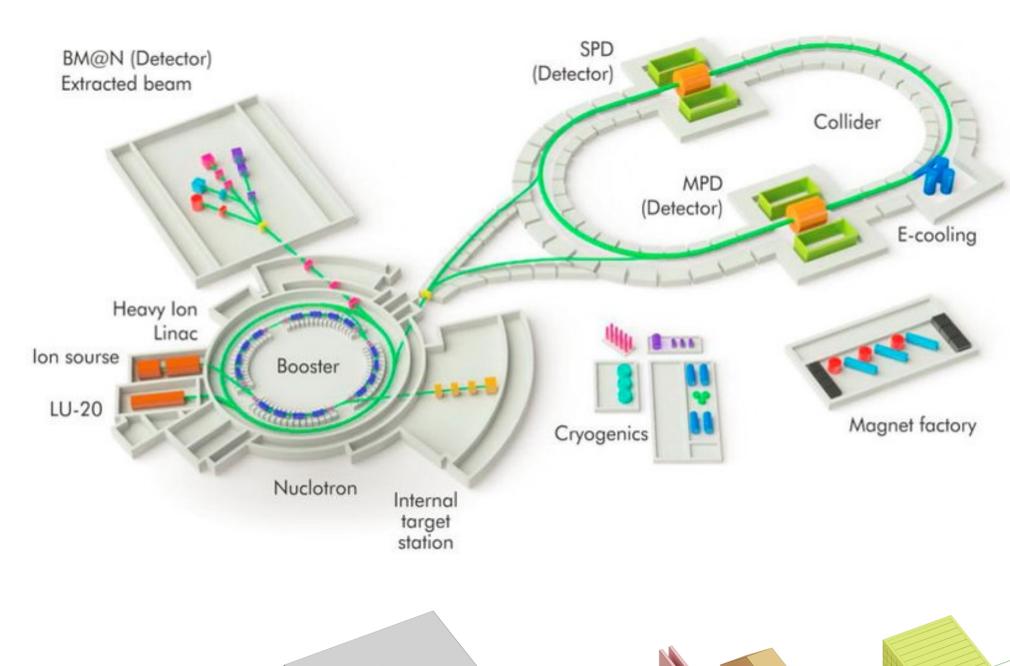


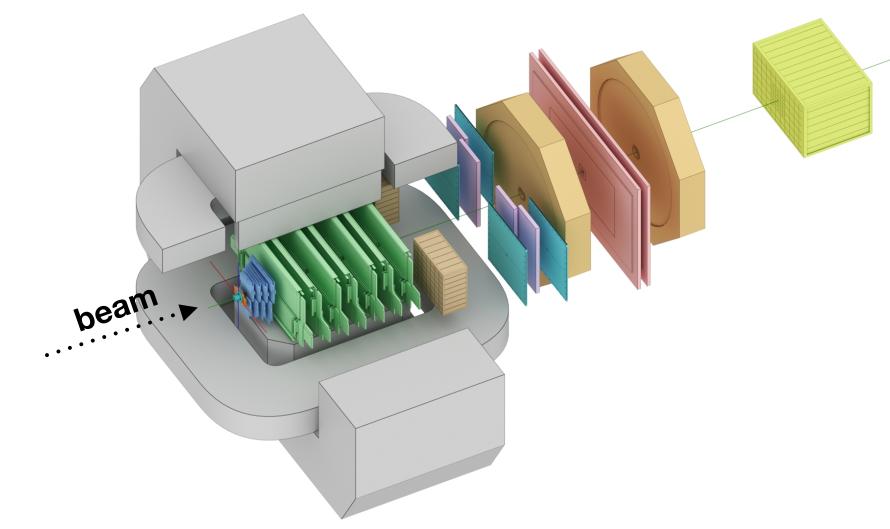


## BM@N experiment

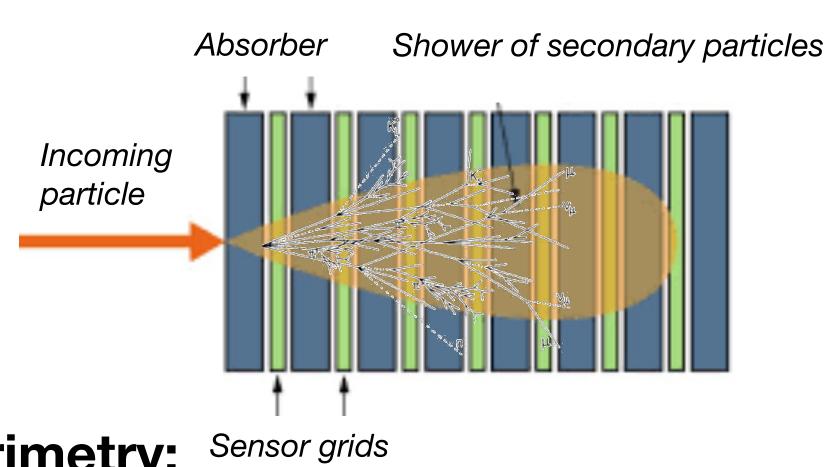
Studies of **B**aryonic **M**atter **at** the **N**uclotron (NICA, JINR Dubna)

- Heavy-Ion beam with energies up to 4A GeV interacts with fixed target
- → investigate the equation-of-state (EOS) of dense nuclear matter which plays a central role for the dynamics of core collapse supernovae and for the stability of neutron stars.
- Azimuthal properties of produced particles important tool for EOS studies
  - •we focus on **neutron** flow and yields





### Neutron detection



**Calorimetry:** 

- Particles interact with matter
- Depends on particle type
  - Neutrons travel on a long distance w/o interaction before producing a shower
- Energy loss transferred to detectable signal
  - Light (e.g. scintillator) -> electric signal
  - Signal collected and acquired
  - Good energy resolution for 5+ GeV n<sup>0</sup>
- In the end: digitized signal
  - Coordinates, energy deposition, time

#### Time-of-Flight (ToF) measurements:

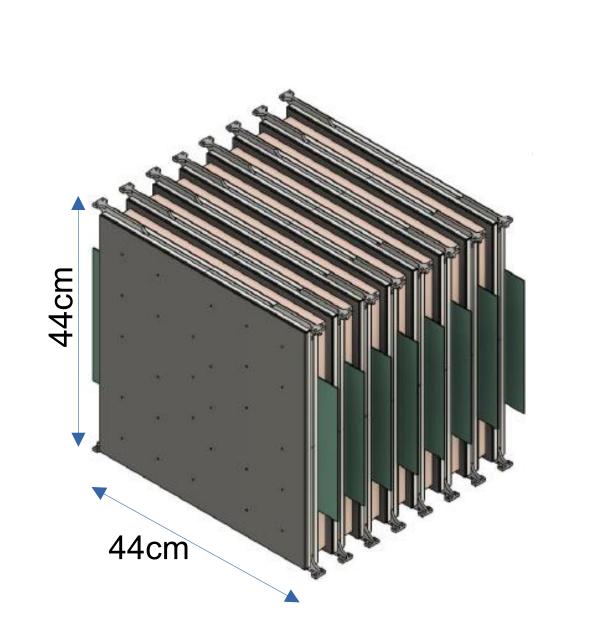
- Time of arrival and distance to the production point gives momentum of a particle
- → With known particle mass (neutron) we can calculate it's energy

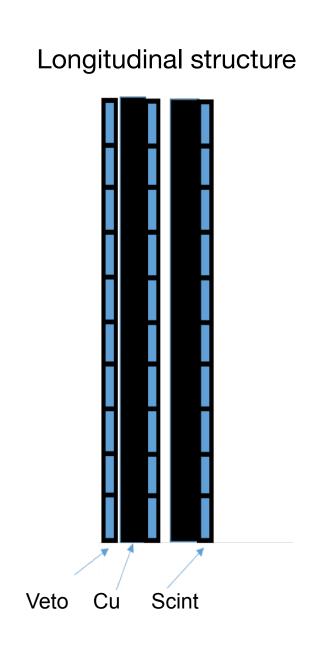
$$E_{ToF} = m_n \left( \frac{1}{\sqrt{1 - (\frac{v}{c})^2}} - 1 \right)$$
 Good energy resolution up to 2-3 GeV n<sup>0</sup>

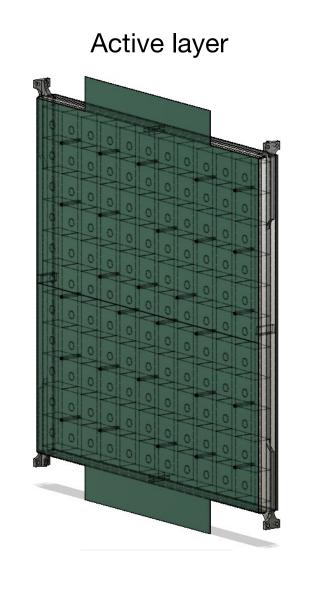
we use calorimeter image to identify neutron and ToF to reconstruct it's energy

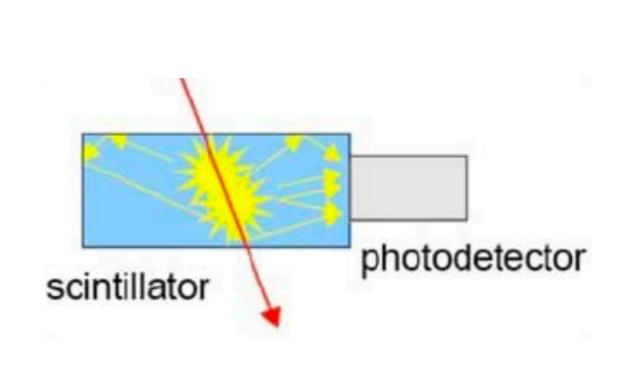
Neutron is one of the most challenging particle types for reconstruction due to relatively low probability to interact with matter. Yield of visible energy is very low at energies < 1GeV

#### Highly granular time-of-flight neutron detector (HGND)





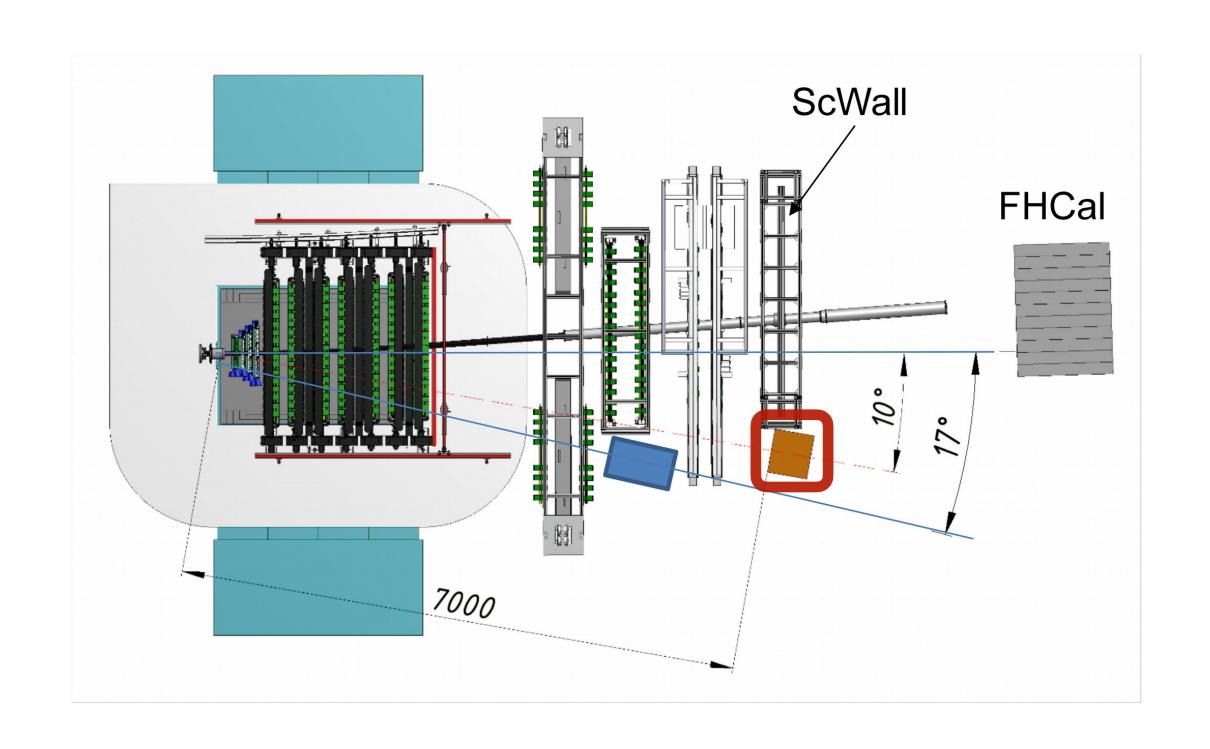


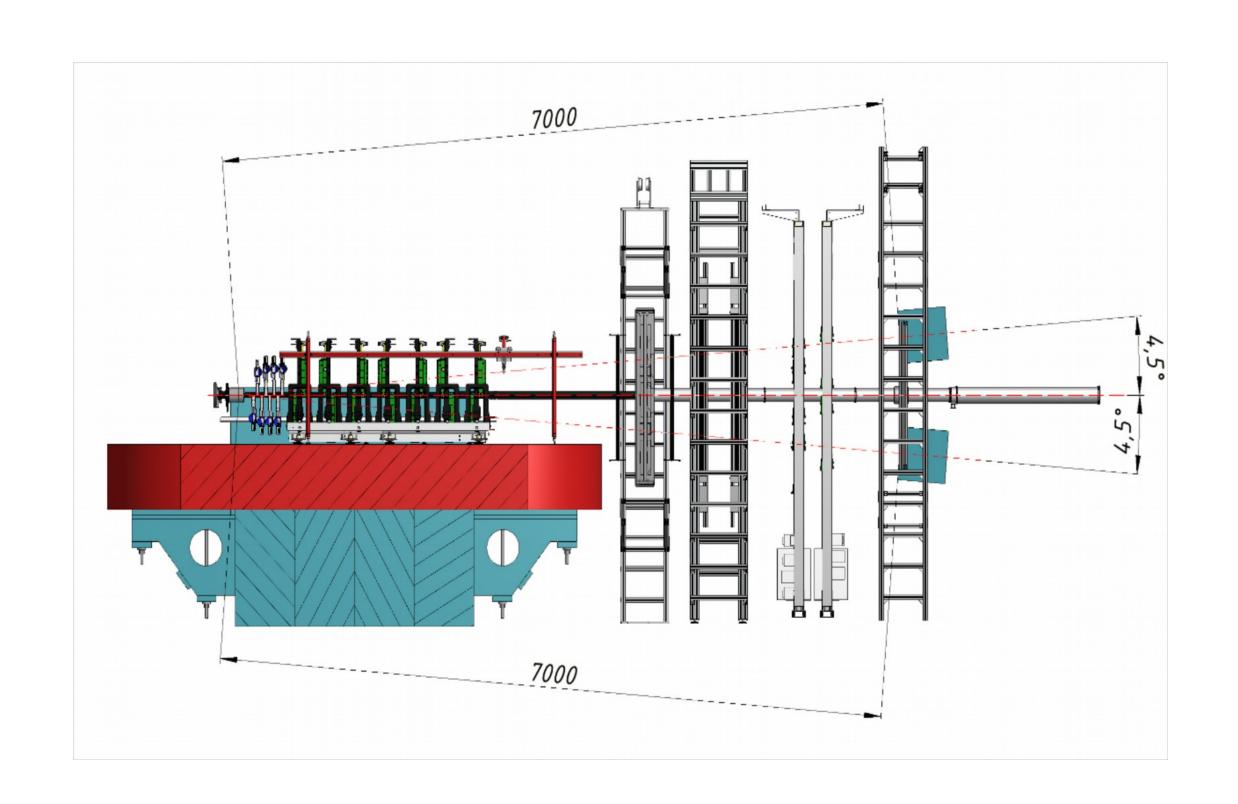


- (2x) 8 layers: 3cm Cu (absorber) + 2.5cm Scintillator + 0.5cm PCB; 1st layer 'veto' before absorber
- →Total length: ~0.5m, ~1.5 λ<sub>in</sub>
- → neutron detection efficiency ~60% @ 1 GeV
- Transverse size: 44x44 cm²
- 11x11 scintillator cell grid

- scintillator cells:
  - •size: 4x4x2.5 cm<sup>3</sup>,
  - total number of cells: 968 (x2)
  - individual readout by SiPM
  - expected time resolution per cell: ~150 ps

### Detector Setup and Simulations





- HGND sub-detectors are located at 10° to the beam axis at ~7m from the target
- Monte-Carlo event simulations with full detector setup model:
- •3 AGeV Bi+Bi DCM-QGSM-SMM model + Geant4 v11.2.0 FTFP\_BERT

•~1M events

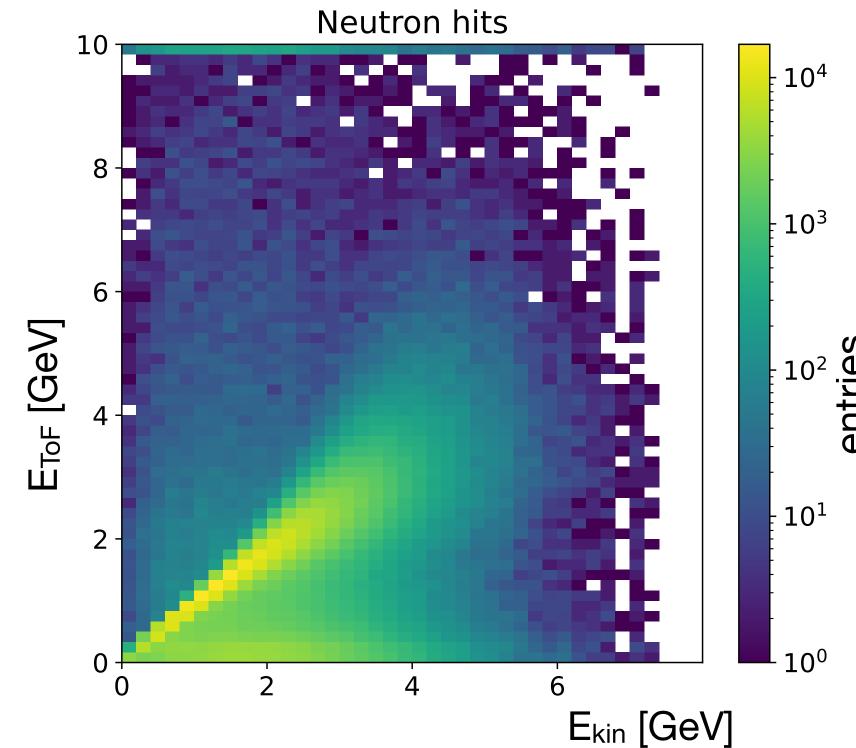
#### Hit Level Information

- $E_{dep} > 3 \text{ MeV} \sim 0.5 \text{ MIP}$
- **ToF energy** for  $n^0$  hypothesis:

$$E_{ToF} = m_n \left( \frac{1}{\sqrt{1 - \beta^2}} - 1 \right)$$

- $t_{hit}+\mathcal{N}(0,\sigma_t=150ps)$
- clip at 10GeV
- Each hit is linked with corresponding surface MC particle
- signal neutrons passed the upstream HGND surface

#### ToF energy for neutron hits



- Underestimation neutron shower development
- Overestimation background contributions in the same hits
  - → More precise labelling is under development

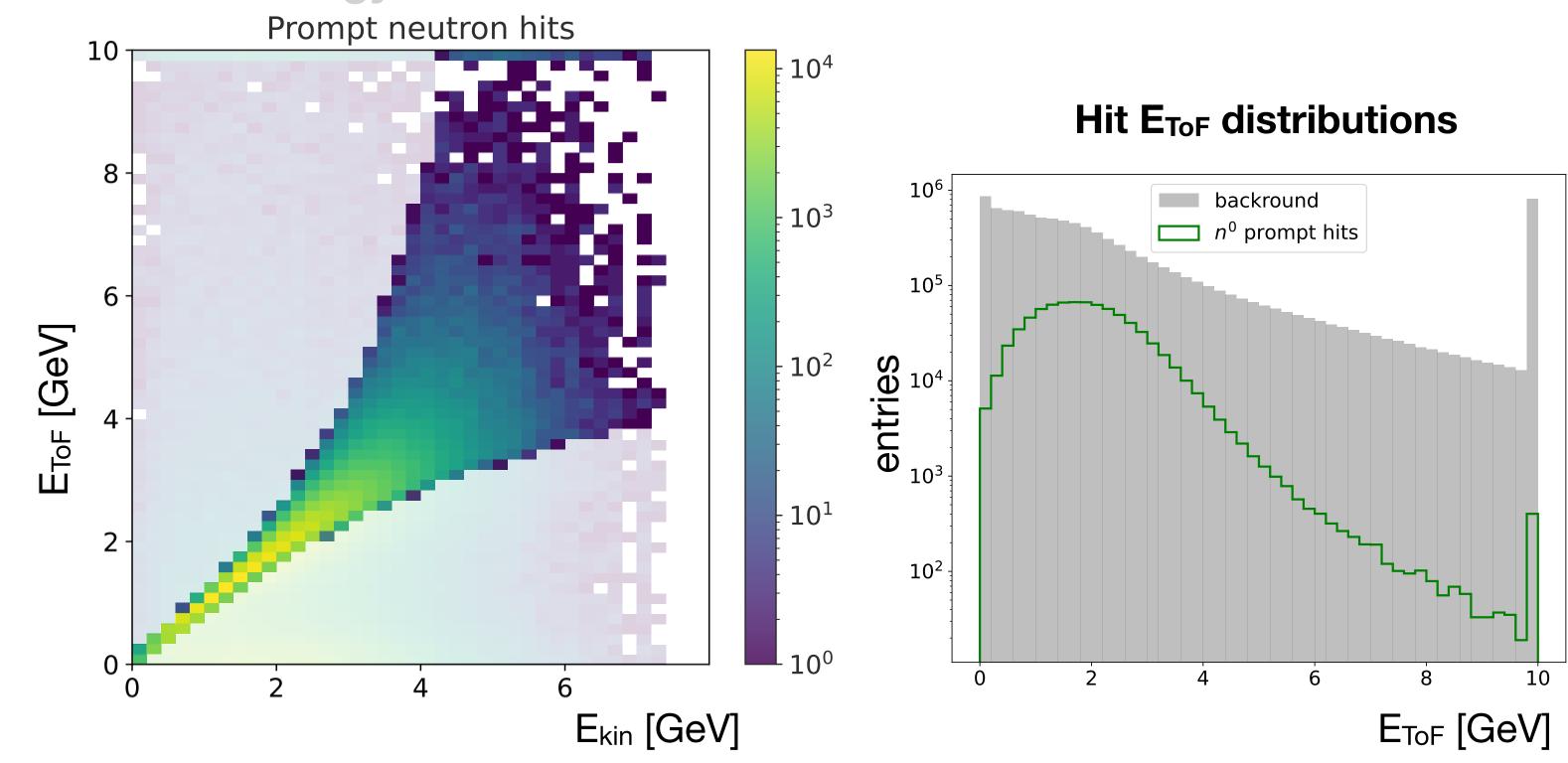
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- $t_{hit} + \mathcal{N}(0, \sigma_t = 150ps)$
- clip at 10GeV
- Each hit is linked with corresponding surface MC particle
- signal neutrons passed the upstream HGND surface
- Prompt neutron deposition
   selected by E<sub>n</sub> ∈ E<sub>ToF</sub>(t<sub>hit</sub> ± 2σ<sub>t</sub>)
  - other hits background

#### **ToF energy for neutron hits**

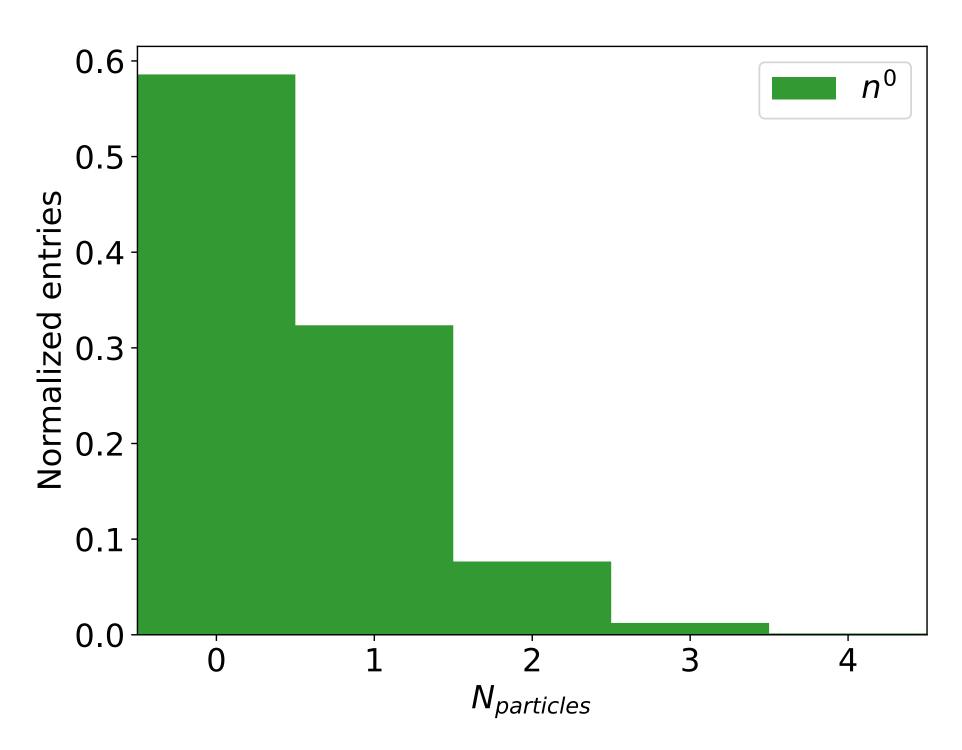


- Underestimation neutron shower development
- Overestimation background contributions in the same hits
  - → More precise labelling is under development

## Neutron Multiplicity

- Multiplicity counts require
  - existence of prompt hit
  - $E_{n0} > 0.1 \text{ GeV}$
  - neutron passed upstream HGND surface
- Distributions normalised to number of events with energy deposition
  - → Neutron detection efficiency is not discussed
- Reconstruction algorithm has to deal with neutron multiplicities > 1

#### Neutron multiplicity per event



# Graph Neural Networks (GNN)

#### Why Graph Neural Networks:

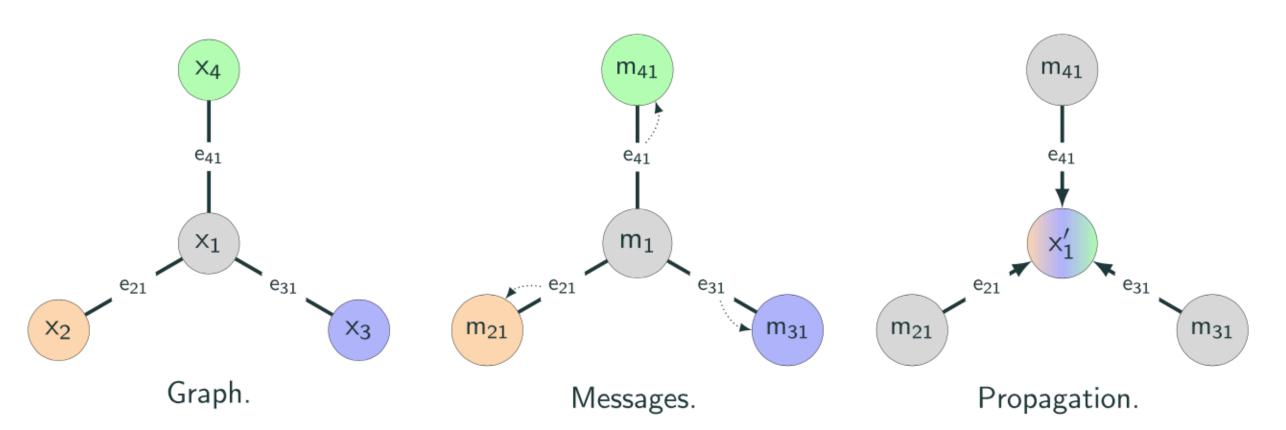
- Natural vector event representation
  - Detector cell hits as graph nodes
- Easily applied to sparse data with variable input size
  - Typically we have signal only in small fraction of sensors
- Captures event structures
- <u>Increasing number</u> of successful implementations in HEP

**HEPML-LivingReview** 

#### Message passing architecture

Key idea:

- Edges propagate information between nodes in a trainable manner to encode local graph structures
- Node embeddings are then aggregated to a problem-specific value, e.g.:
  - Graph/hit class "probability" signal/background
  - Target value neutron energy



J. Gilmer et al., "Neural message passing for quantum chemistry," 2017.

### Graph Construction

- Nodes hits. Observables per hit:
  - hit coordinates: layer, row, column
  - $E_{dep} > 3 \text{ MeV} \sim 0.5 \text{ MIP}$
  - hit time +  $\mathcal{N}(0,\sigma = 150 \text{ps})$
  - E<sub>ToF</sub>
- Edges
  - Predefined clustering:
    - radius graph. R = 3.6 cells
    - time window 1.5 ns
    - cluster isolated subgraph

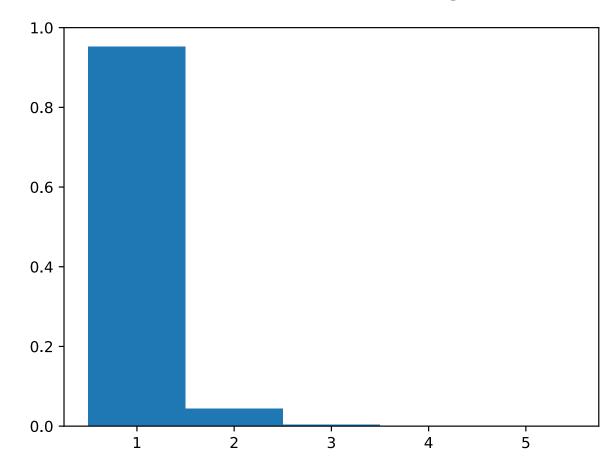
#### Labeling for further training

- Cluster level:
  - Signal cluster contains at least 1 prompt neutron hit
  - Energy of fastest neutron with prompt hit in a cluster

## Clustering

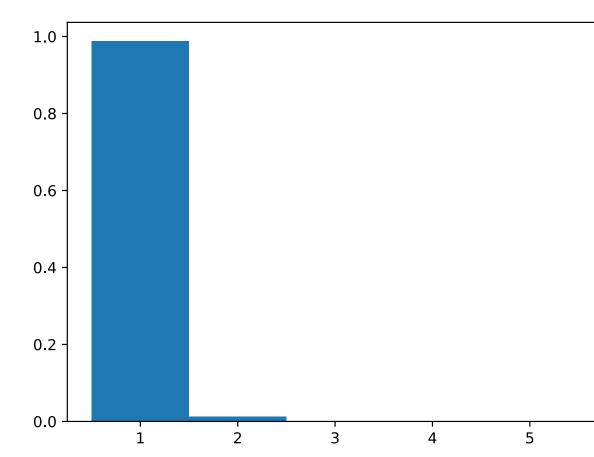
- Soft spatial rule.
  - search radius R = 3.6 cells
- ~strict temporal rule
  - time window 1.5 ns
- ⇒ first guess parameters, to be optimised, included in GNN



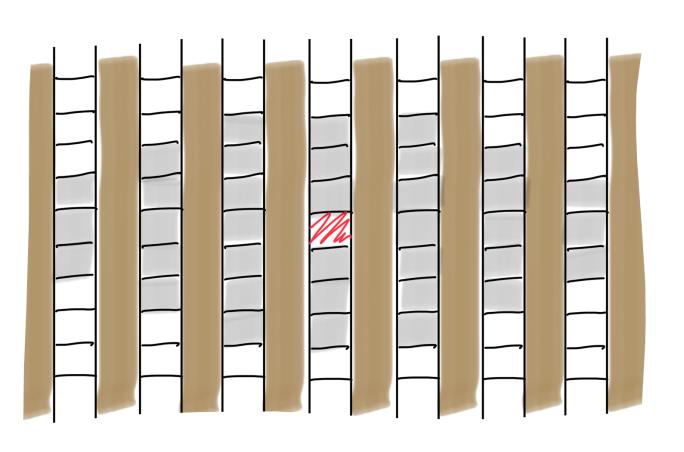


isolation >95%

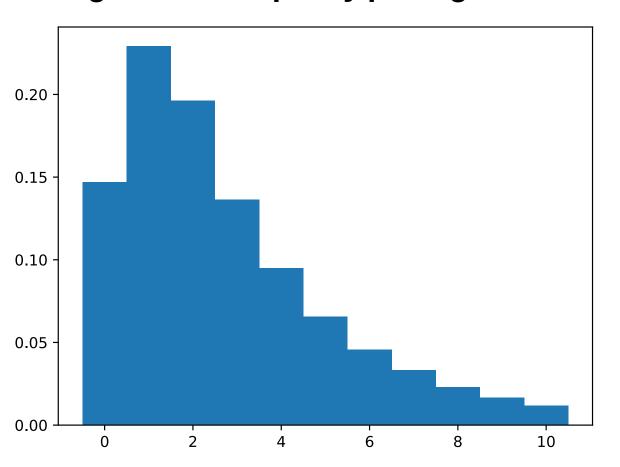
Cluster multiplicity per prompt n<sup>0</sup>



 prompt n<sup>0</sup> splits in secondary clusters at level of < 2%</li>



Background multiplicity per signal cluster



significant background
 contribution => make use of ML

#### GNN Model

#### **Training objective**

- Cluster level:
  - predict neutron class score
  - reconstruct expected neutron energy

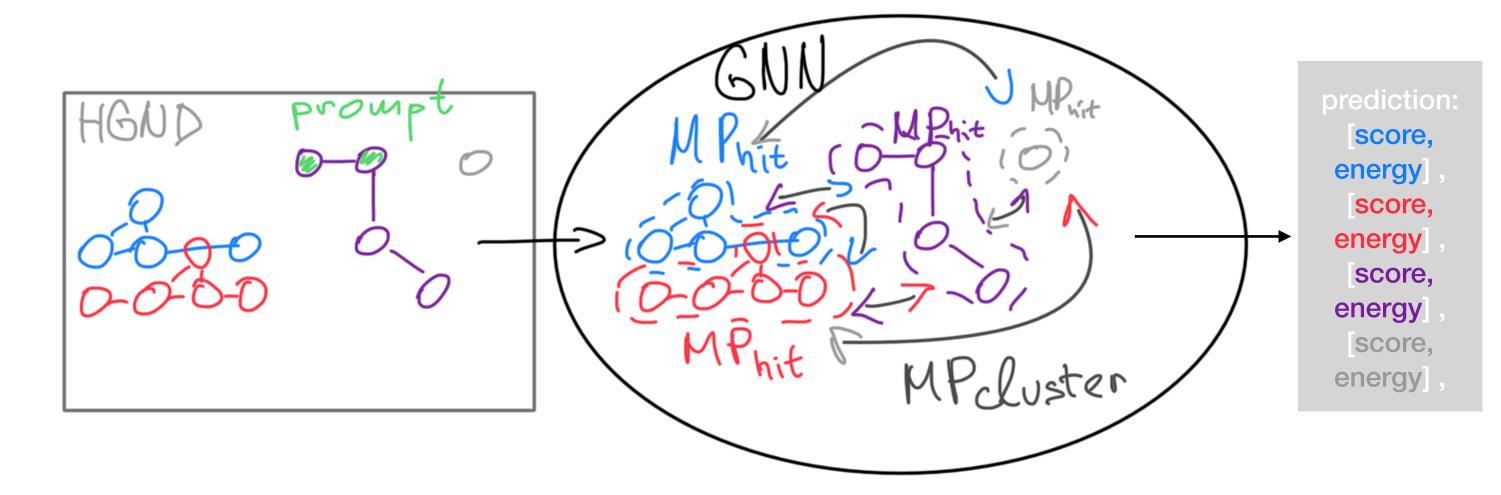
#### Loss function

- Binary Cross Entropy for classification
- Mean Squared Error for energy regression

$$Loss = BCE_{cluster} + MSE_{cluster}$$

#### Heteregenious GNN architecture

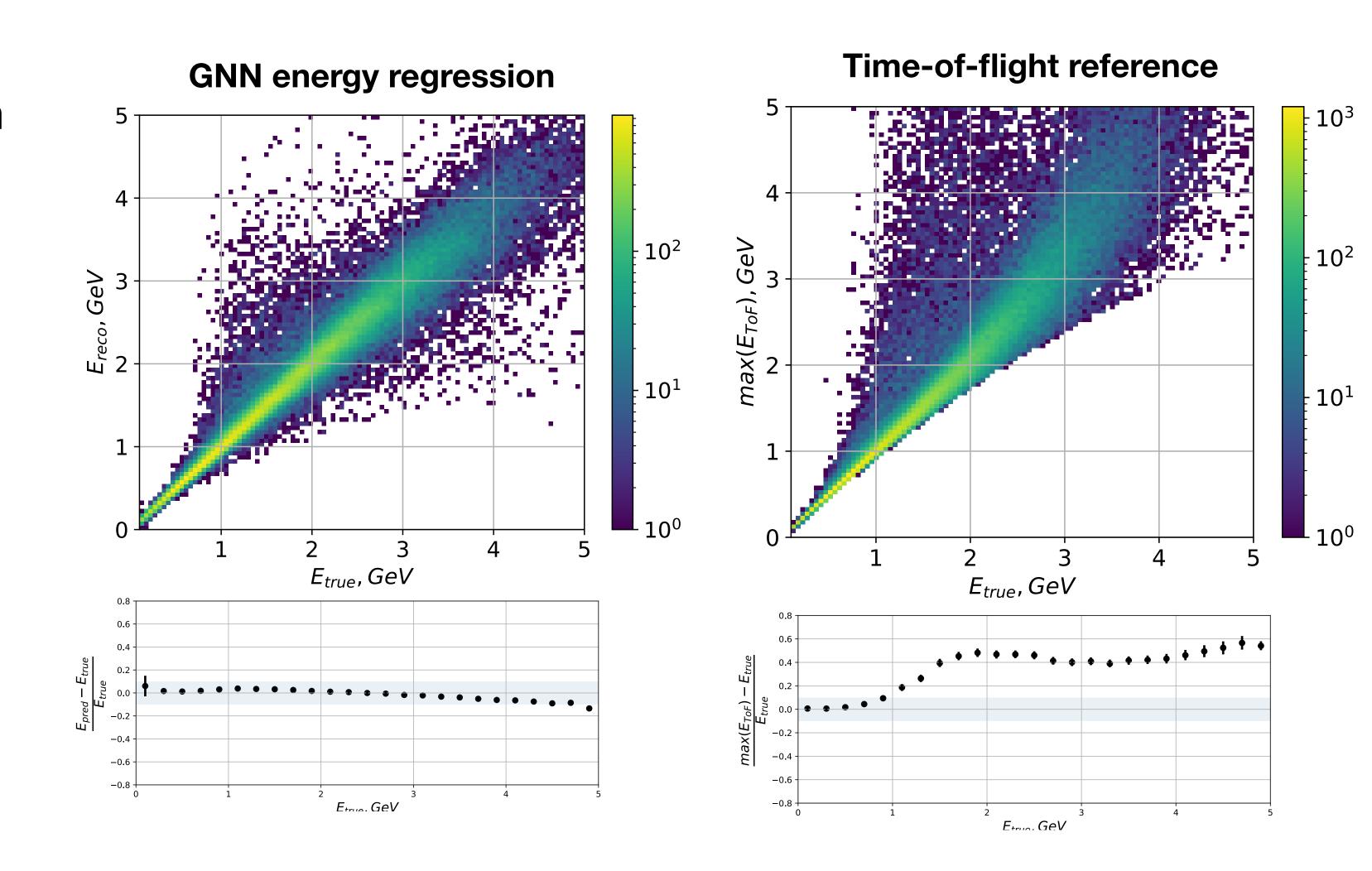
- MP<sub>hit</sub>. EdgeConv+GraphSage layers hit -> hit within clusters
- GraphConv layer to aggregate hit nodes -> virtual cluster node
- MP<sub>cluster</sub>. FC EdgeConv layer. cluster -> cluster
- Cluster output: class score, predicted energy





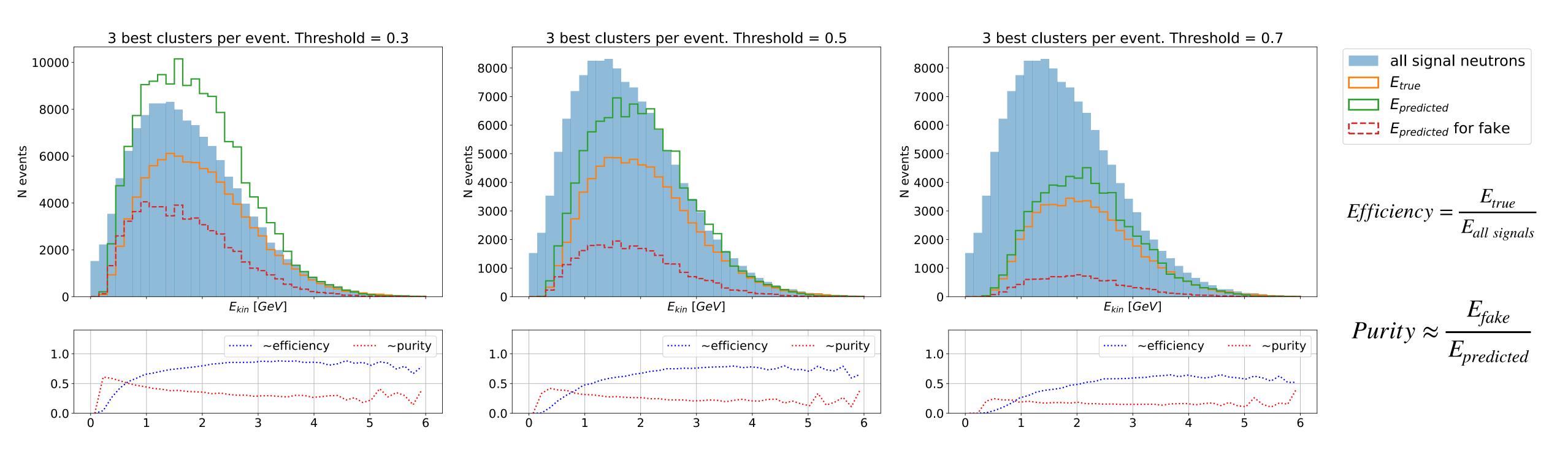
### Cluster Reconstruction Performance

- Overall good cluster classification performance. ROCAUC≈0.96
- Energy resolution ≈13% estimated on integrated relative error for true neutron clusters
  - Linearity within 10% for the most part of energy spectrum. Model compensates ToF overestimation



# Neutron Energy Spectra

#### Examples of resulting neutron energy spectra

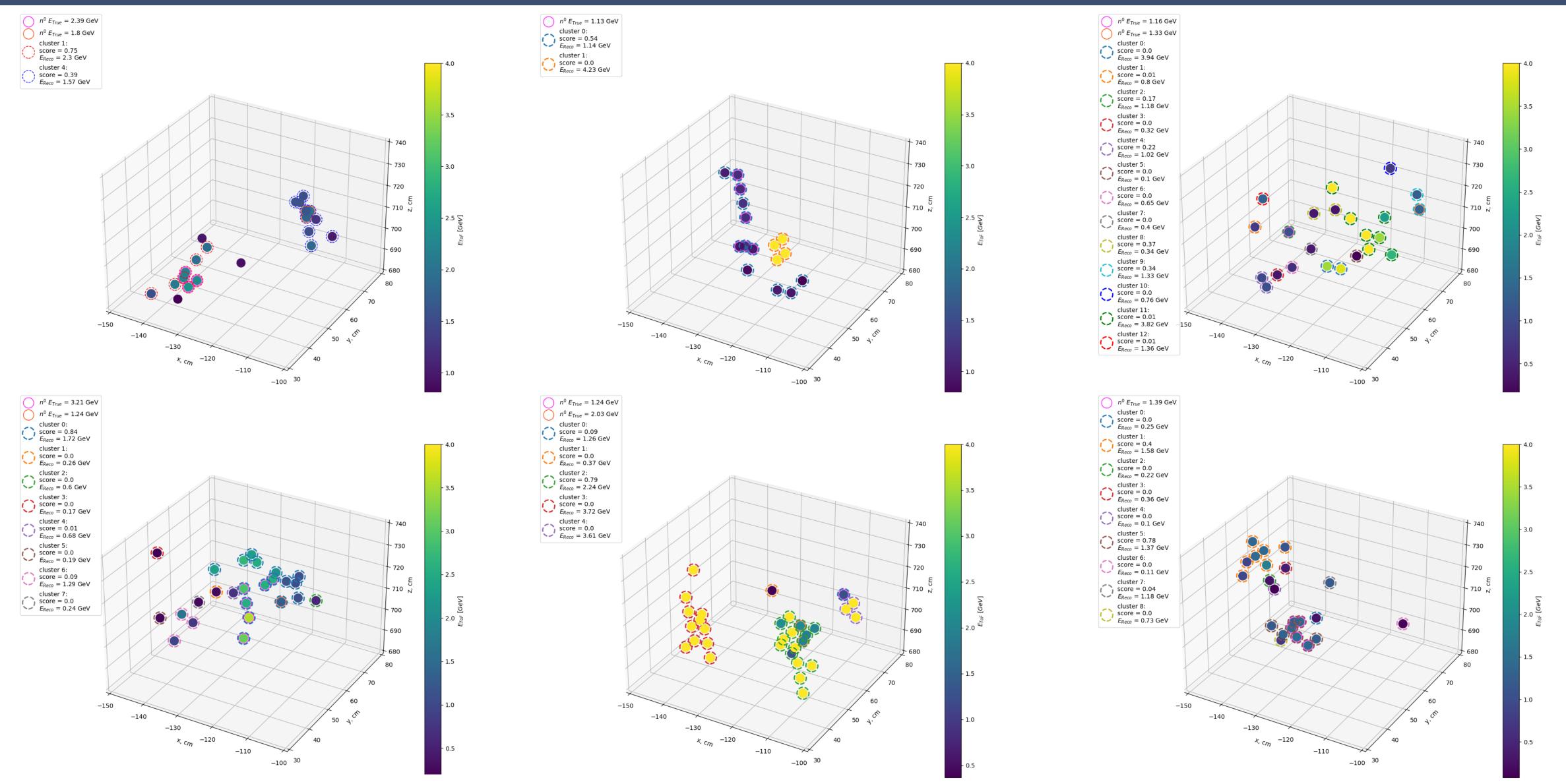


- Energy dependent reconstruction performance
- Trade between purity and efficiency
- Optimal reconstruction scenario will be optimised on end-to-end physics performance simulation

### Summary & Outlook

- Graph Neural Network-based neutron reconstruction algorithm in the highly granular time-offlight detector is presented
- Further steps for GNN model development:
- Implementation of differentiable clustering
- Detailed study of background contributions
- → Model architecture development
- Final physics performance and optimal reconstruction procedure will be defined using parametrised simulations
- → Generative model is under development

# Event Displays



## Physics performance toy-test

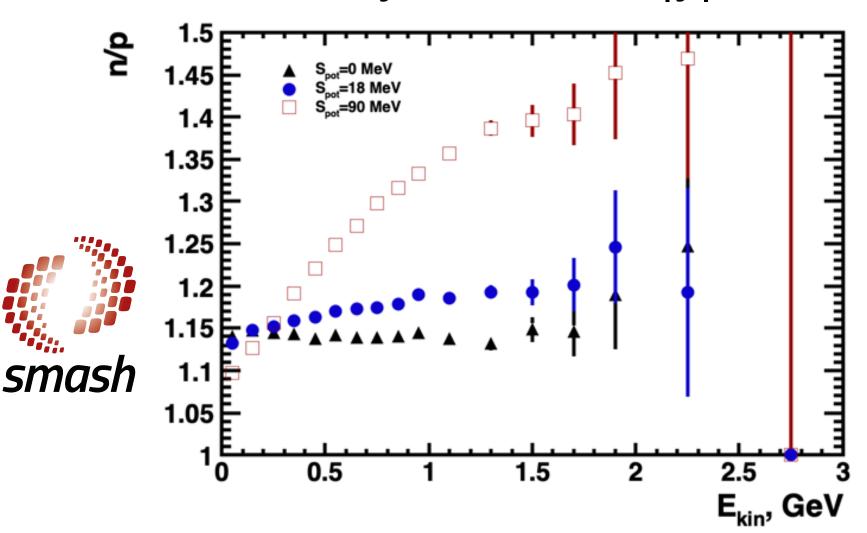
Resample test HGND dataset to get different neutron spectra:

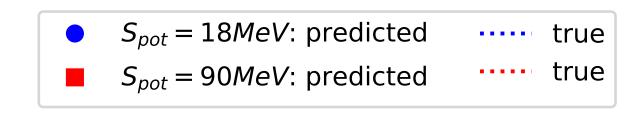
#### SMASH simulations Xe+Cs @ 3.8A GeV:

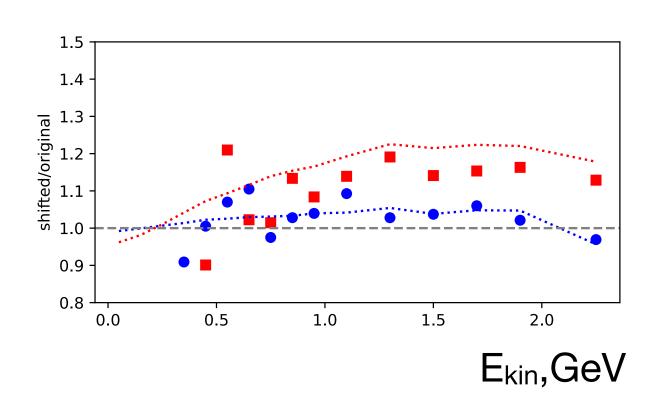
500k events for 3 symmetry potential (S<sub>pot</sub>) values

### Neutron to proton ratio vs $E_{kin}$ . All reaction particles.

Centrality selection: |y| < 0.5







# Physics performance toy-test

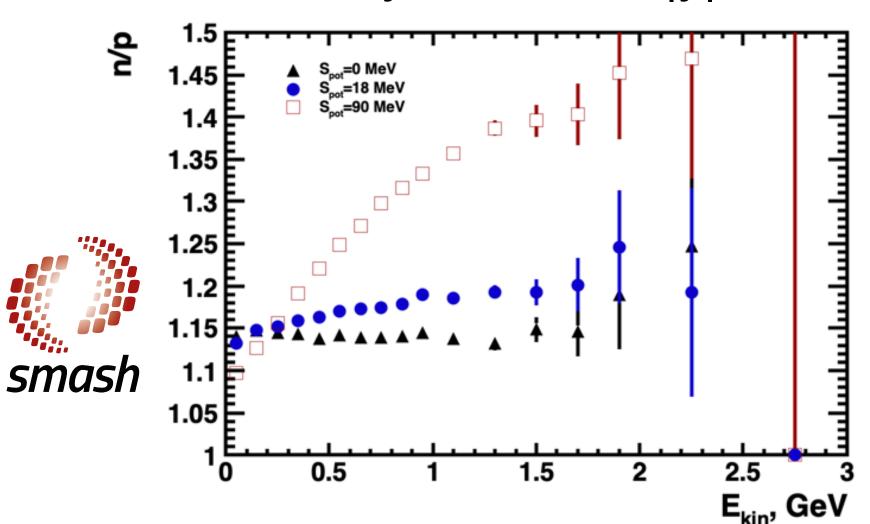
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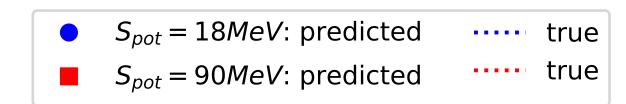
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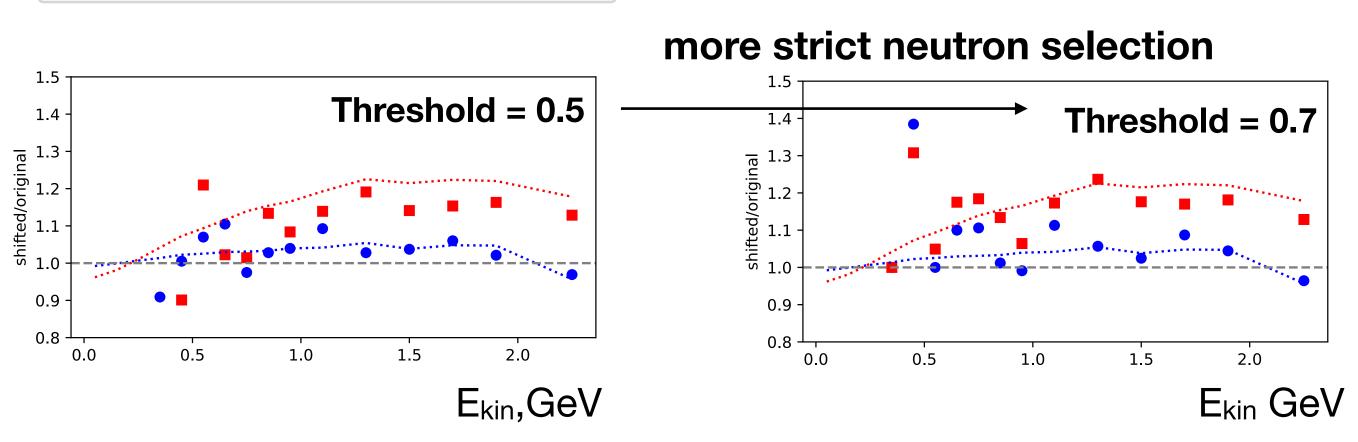
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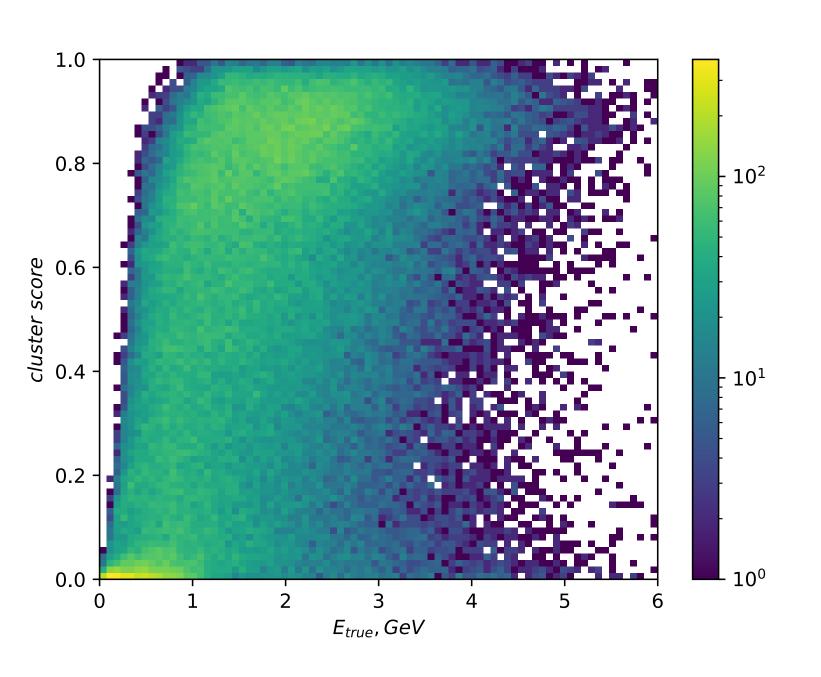


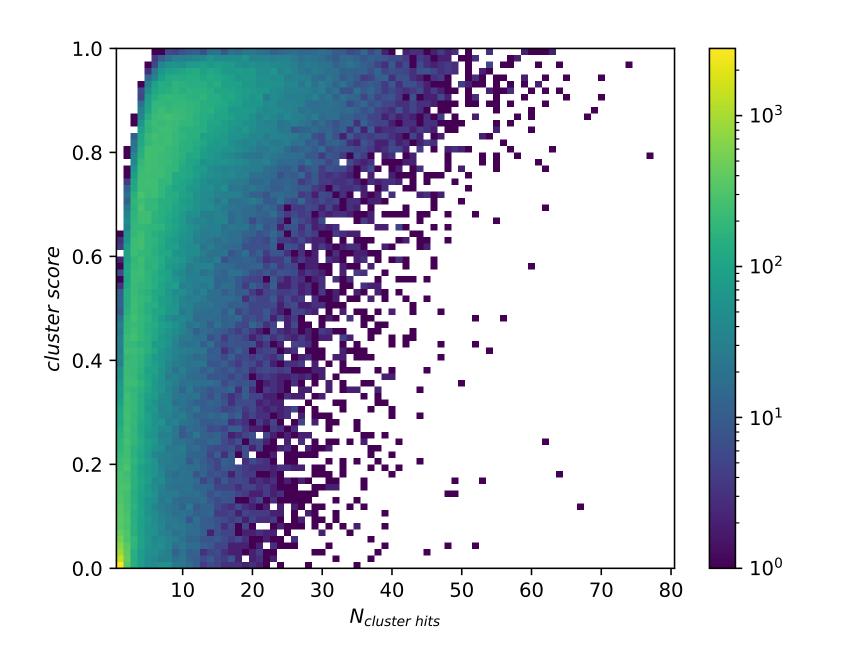


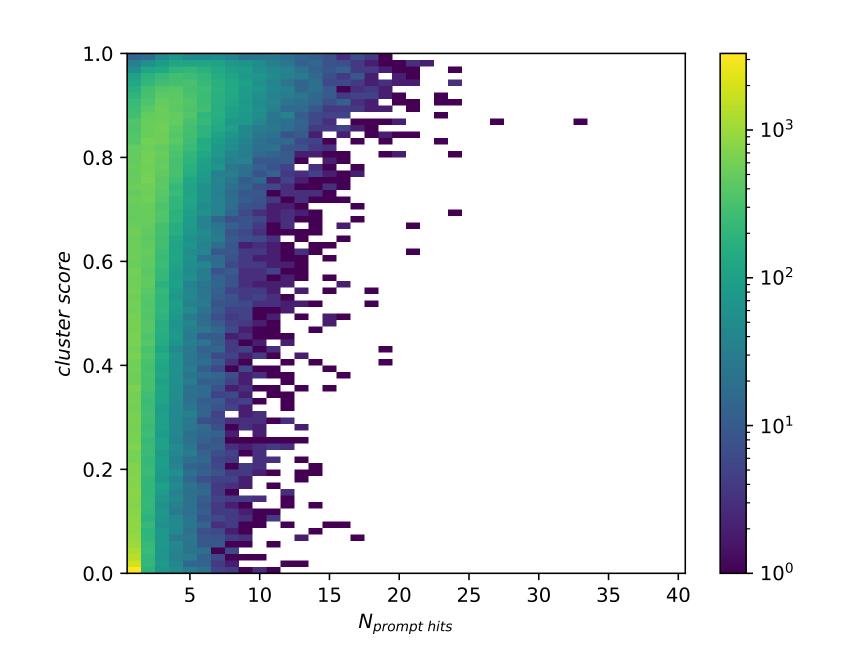


- Qualitatively, it's possible to separate different symmetry potential at ~1GeV+ with relatively strict selection criteria
- Detailed understanding of background contributions is essential

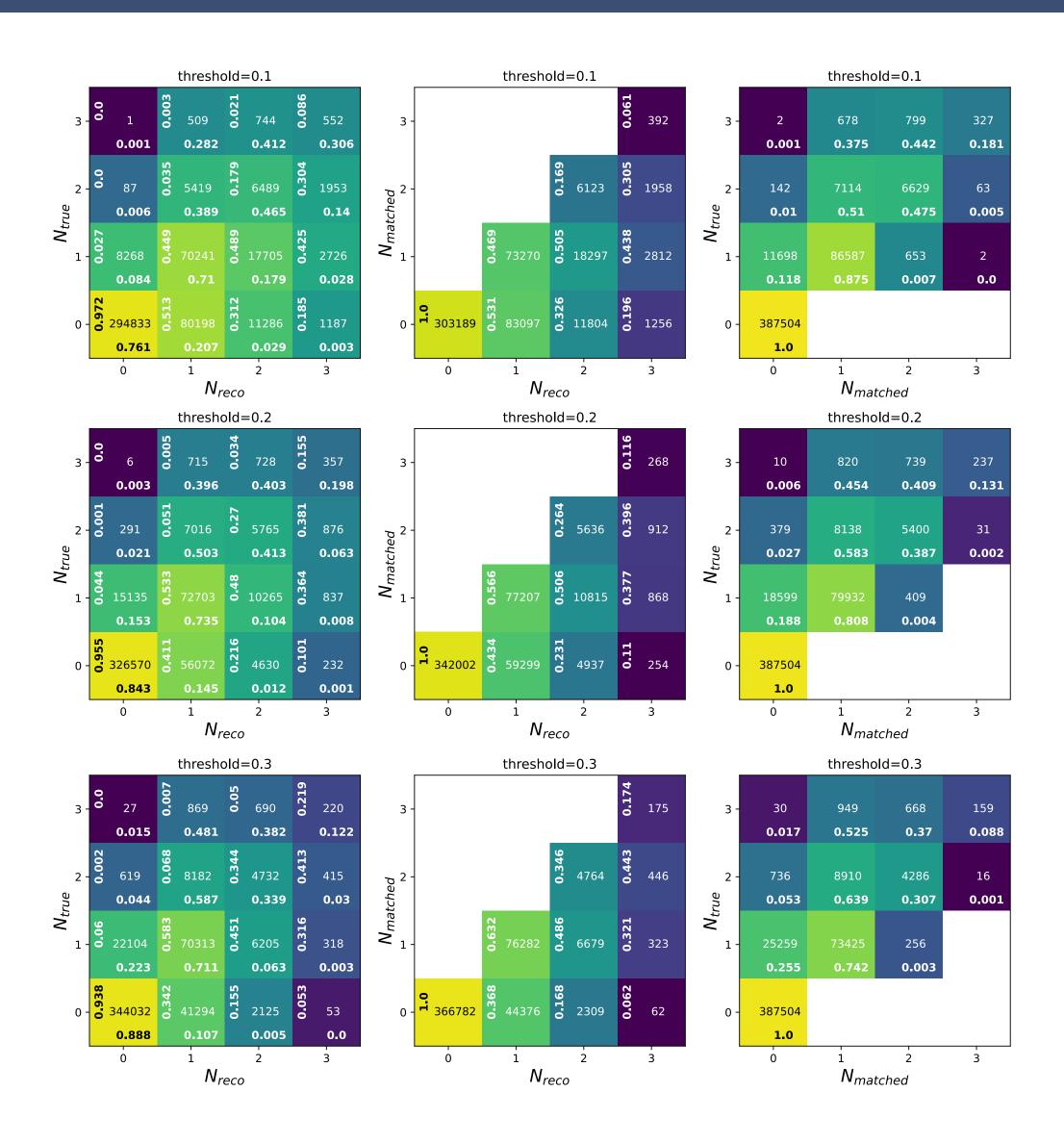
### Score for different clusters

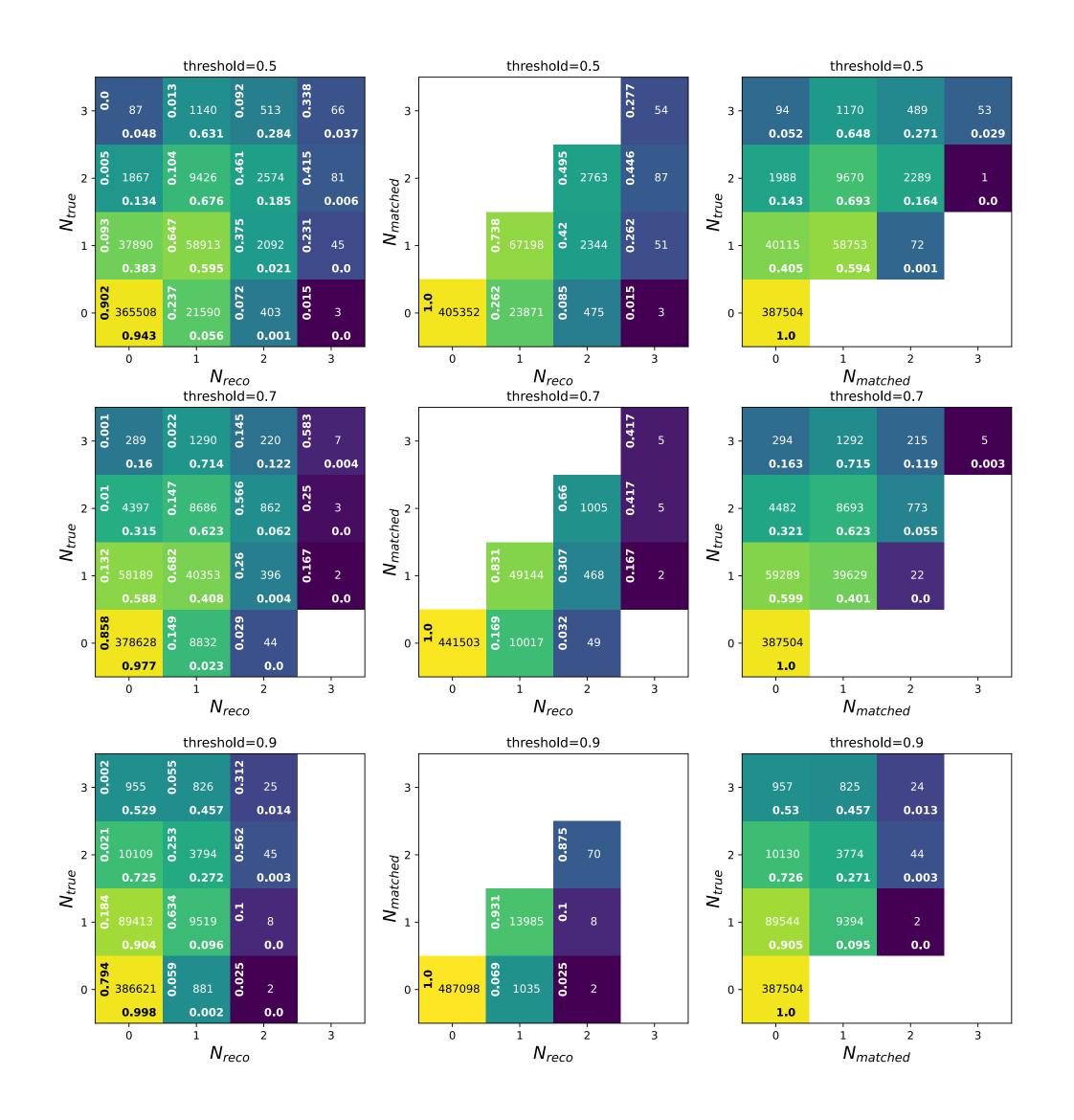




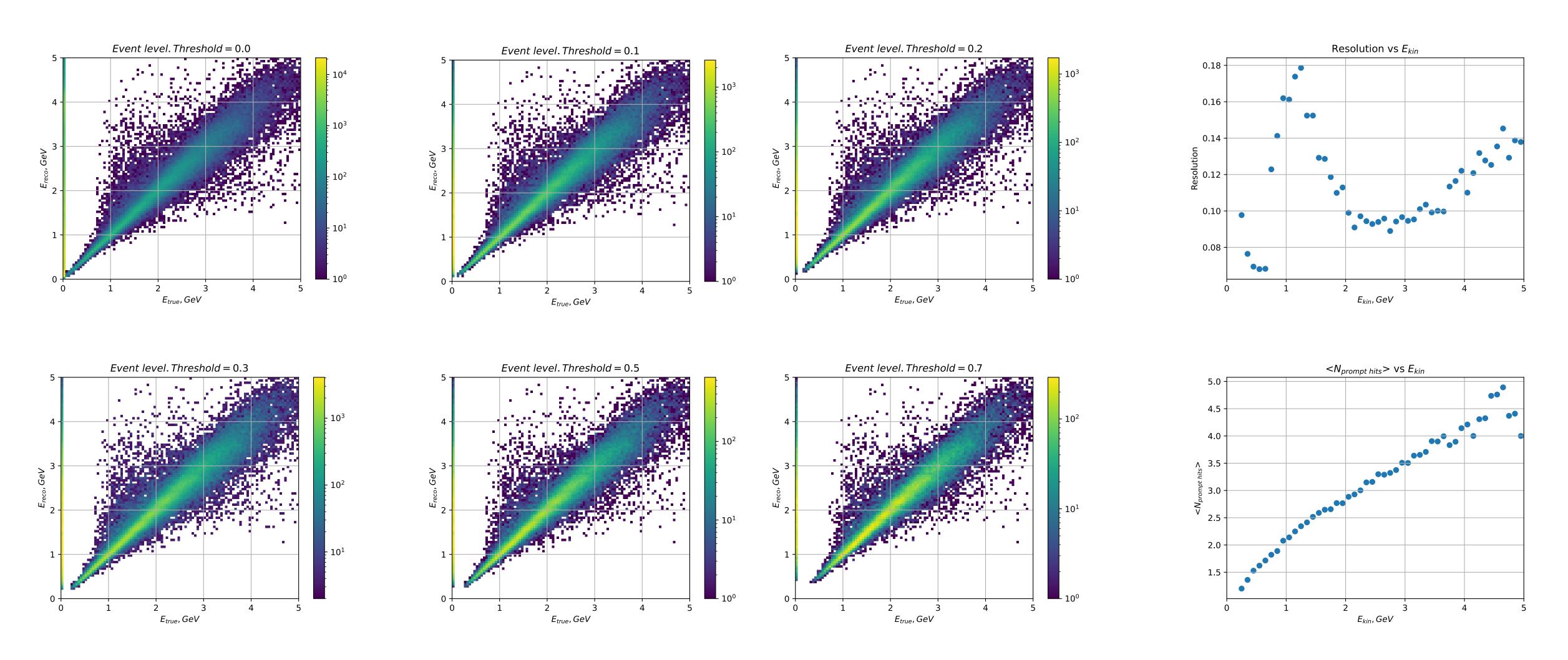


### Multiplicity Reconstruction

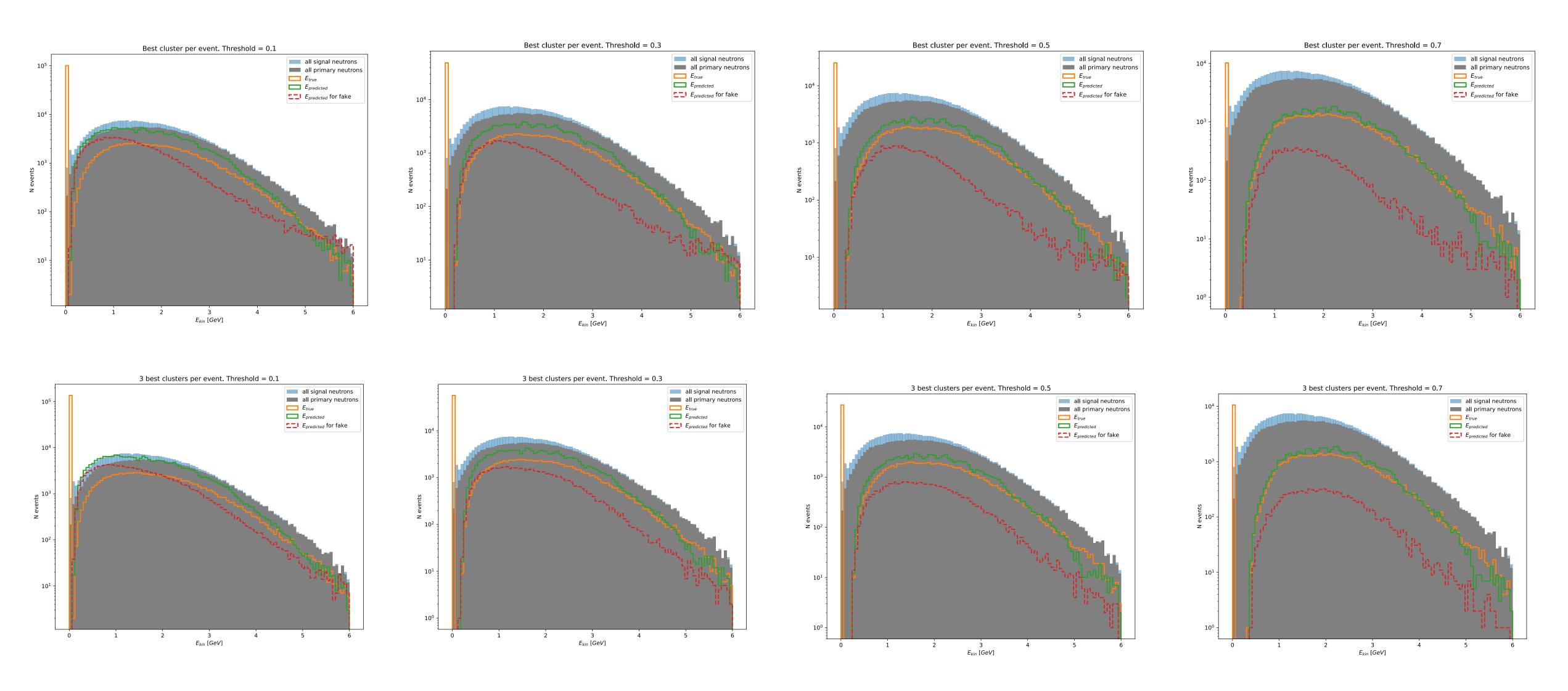




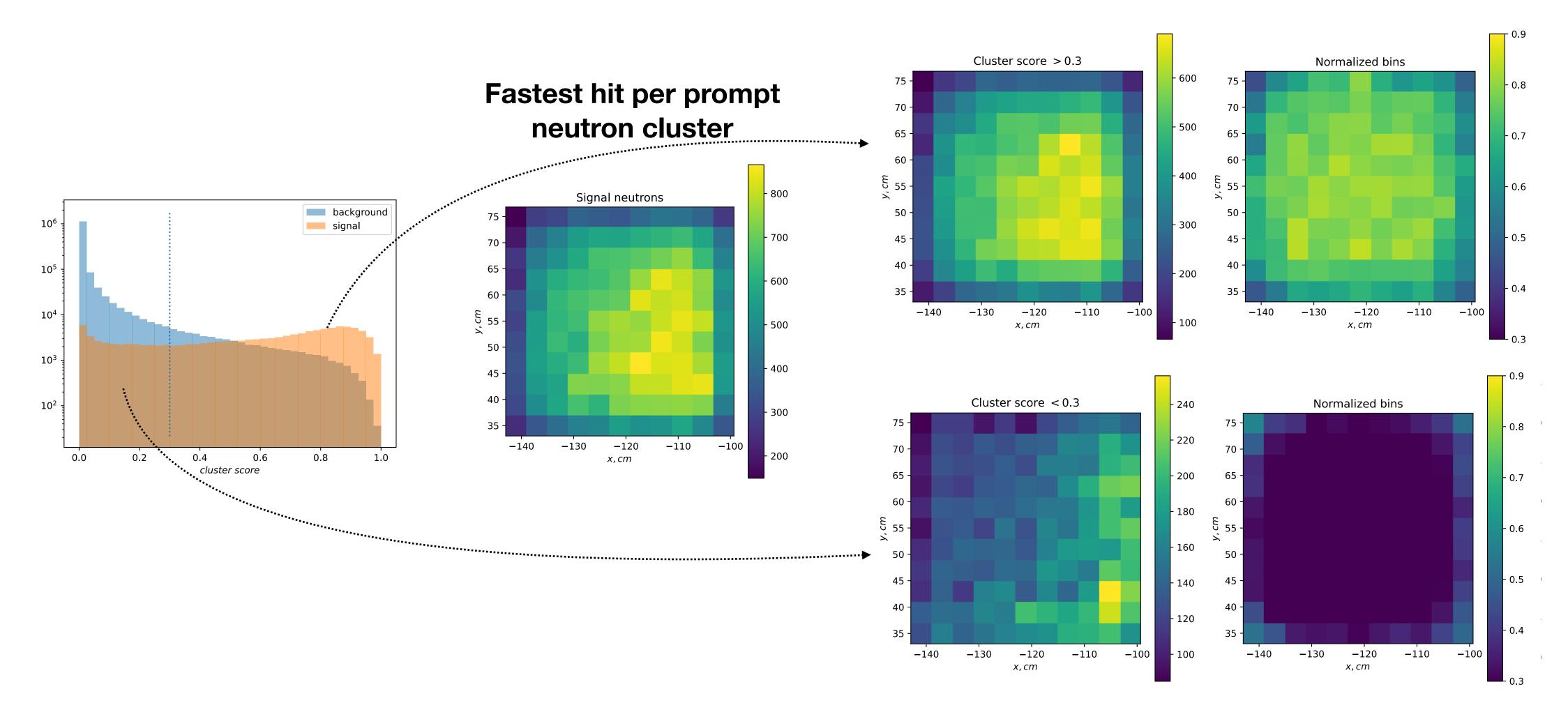
# Energy Reconstruction



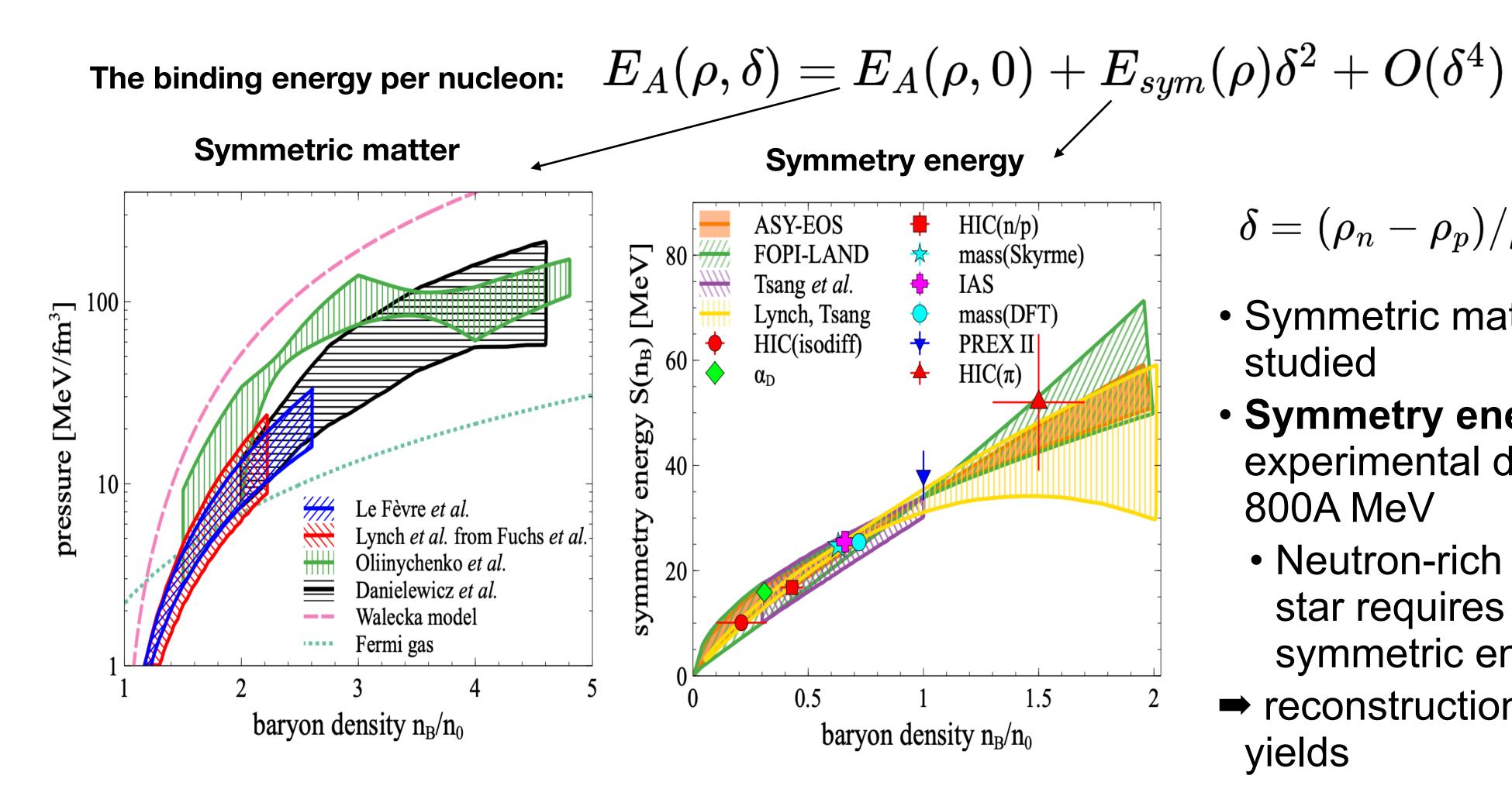
### Spectra Reconstruction



### Efficiency in the XY-plane



# EOS for high baryon density matter



$$\delta = (
ho_n - 
ho_p)/
ho$$
 - Isospin asymmetry

- Symmetric matter is extensively studied
- Symmetry energy contribution lacks experimental data at energies above 800A MeV
  - Neutron-rich matter like a neutron star requires measurement of symmetric energy parameters
- → reconstruction of neutron flow and yields

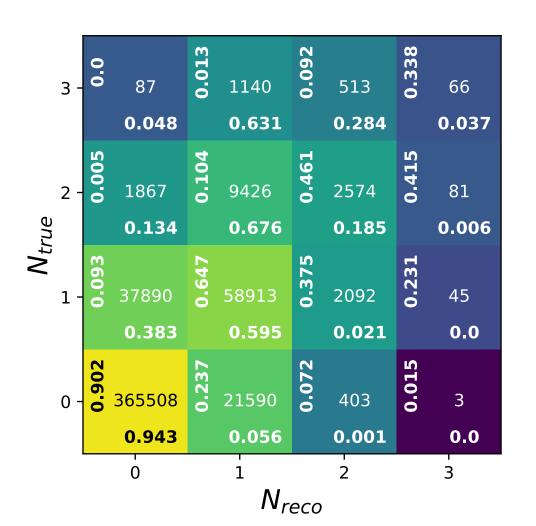
A. Sorensen et. al., Prog.Part.Nucl.Phys. 134 (2024) 104080

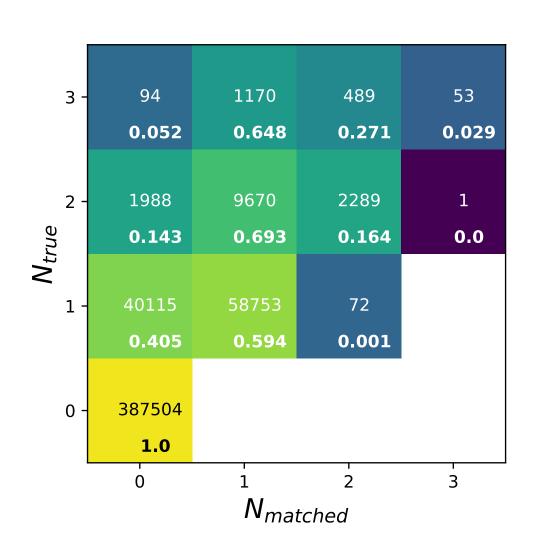
## Multiplicity Reconstruction

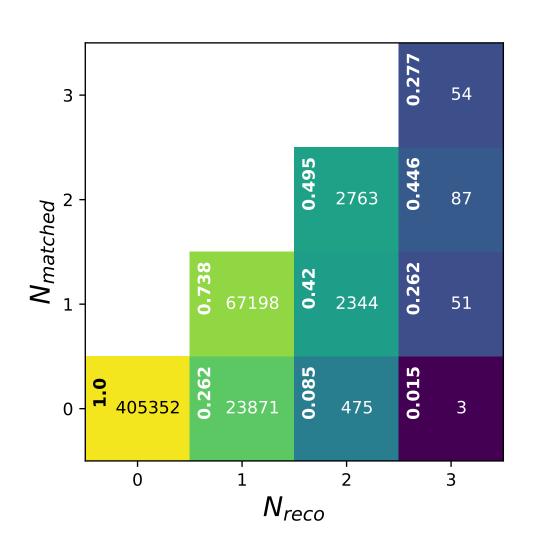
- Test dataset
  - 502198 events
- Selection:
  - E<sub>true</sub> > 0.1 GeV
  - E<sub>reco</sub> > 0.1 GeV

- 4 multiplicity classes:
  - [0, 1, 2, 3 and more]
- Fixed cluster score threshold = 0.5
- horizontal normalisation is related to efficiency
  - efficiency decreases for higher true multiplicities and higher score threshold
- vertical normalisation is related to purity
  - high score threshold allows to separate up to 3 'good' neutron clusters

Matched neutron - reconstructed cluster corresponds to MC-truth neutron cluster.

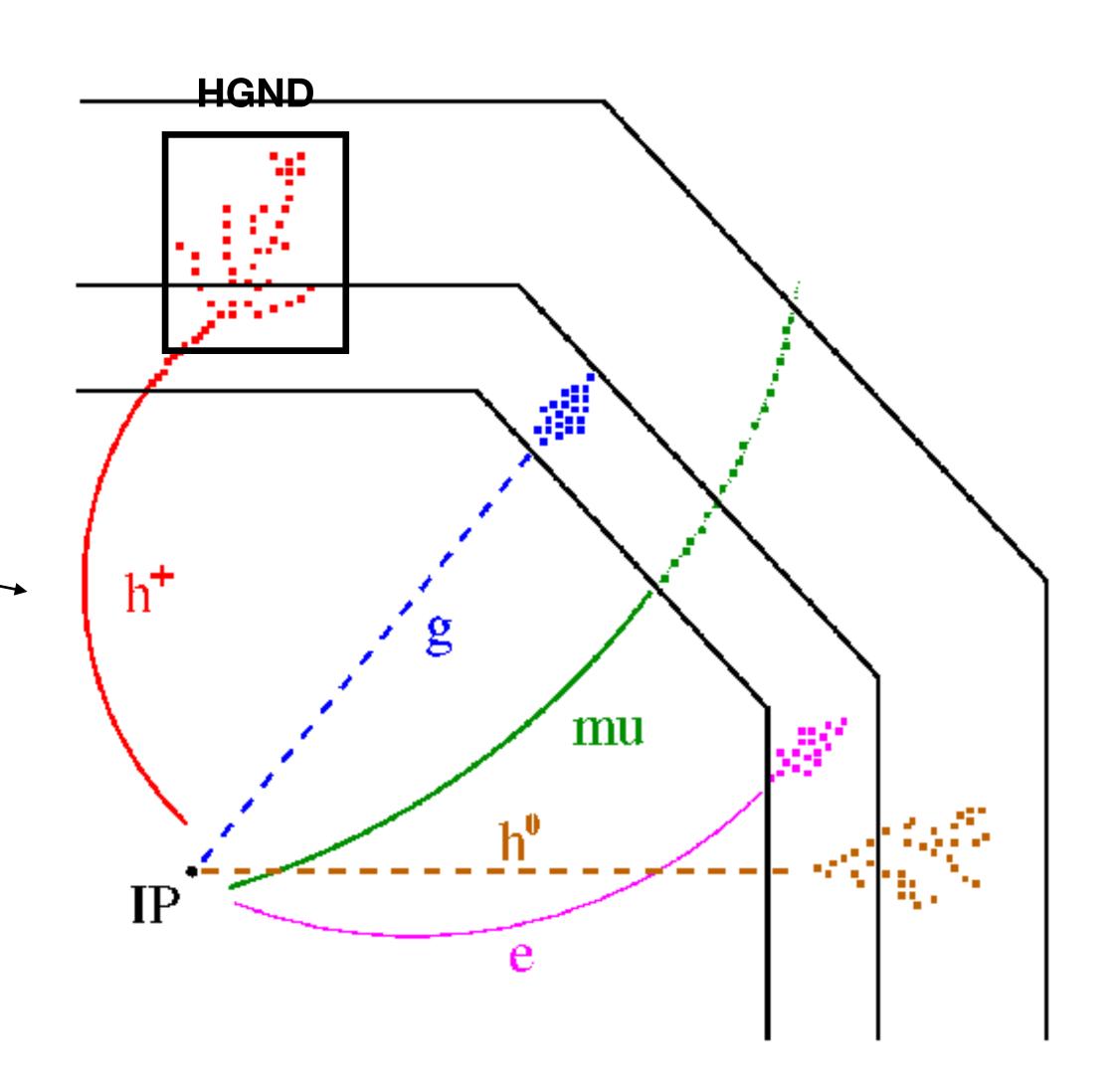






# Tracking Outlook

- BM@N has a dedicated tracking system to reconstruct charged particles
- Tracks may be projected to the HGND upstream surface
- → Additional source of information to reject clusters from charged particles

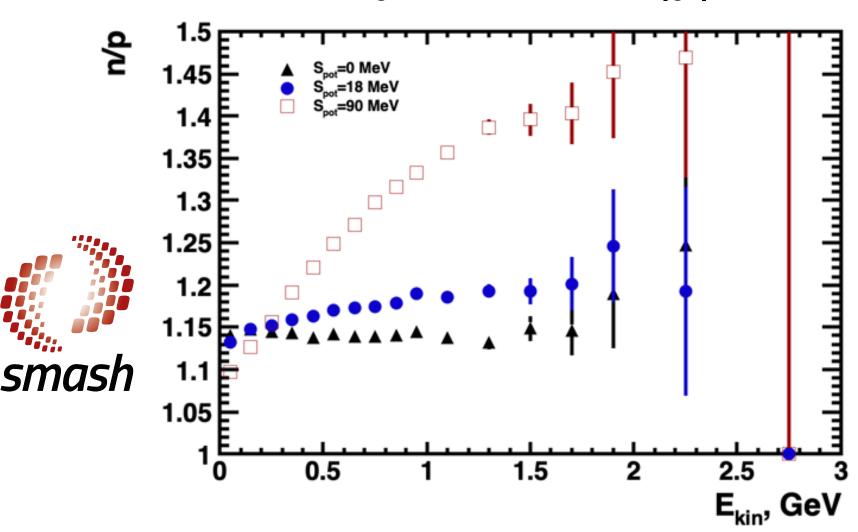


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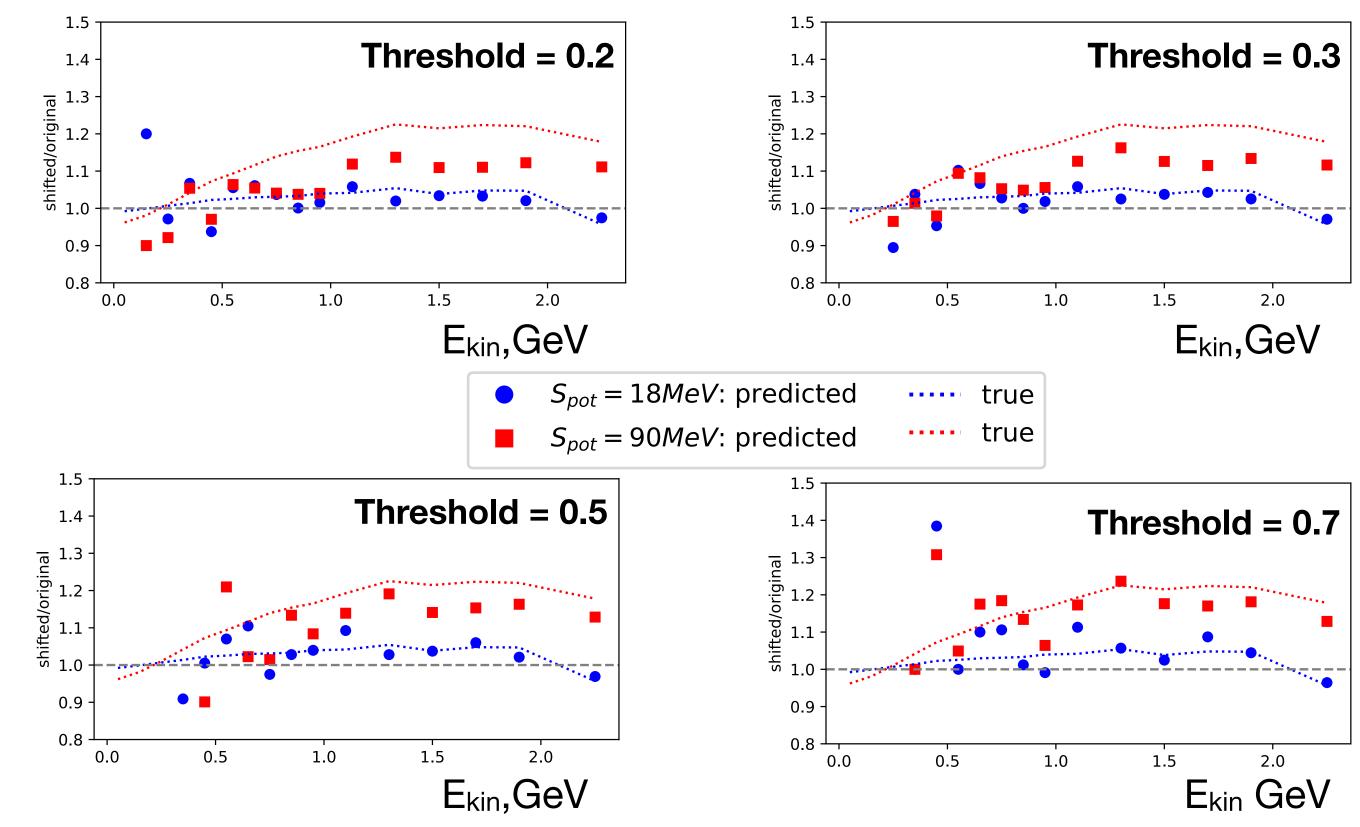
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Centrality selection: |y| < 0.5



Resample test HGND dataset to get different neutron spectra:



- Qualitatively, it's possible to separate different symmetry potential at ~1GeV+ with relatively strict selection criteria
- Detailed understanding of background contributions is essential

#### Event Level Performance

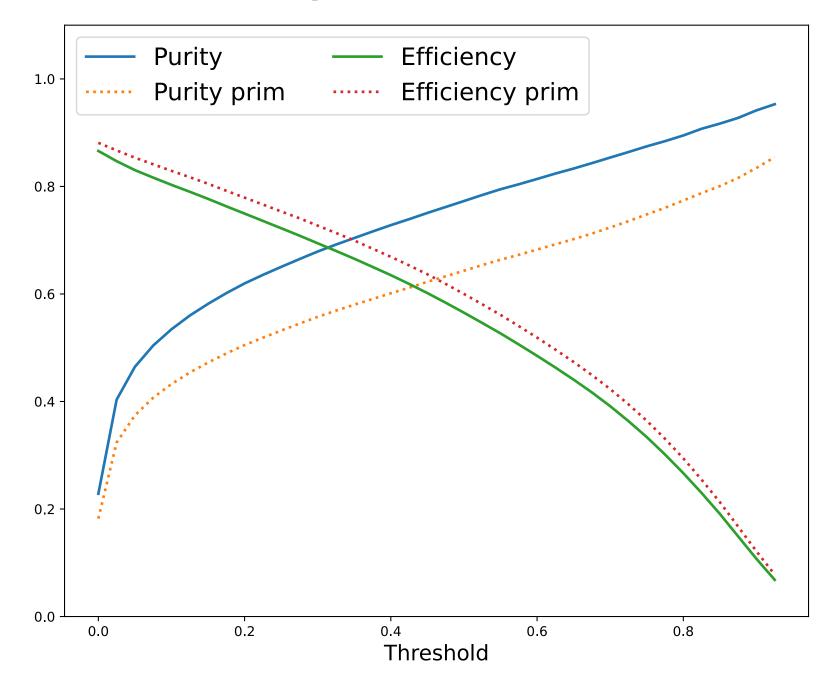
#### Single neutron reconstruction approach:

- Select best cluster in event by cluster score
- Varying threshold for event score and calculate neutron reconstruction efficiency and purity

$$Purity = \frac{N_{reco\ true}}{N_{reco\ all}}$$

$$Efficiency = \frac{N_{reco\ true}}{N_{neutrons}}$$

#### **Event classification performance vs score threshold**



• At selected threshold 0.5 Efficiency ≈ 0.57, Purity ≈ 0.77

#### Motivation

Measurements of neutron flow and yields require reconstruction of neutrons

Neutron reconstruction task:

- Identify neutrons produced in reaction in presence of background
  - **→** use of high granularity
- Reconstruct neutron kinematics:
  - Kinetic energy time-of-flight (ToF) method
- Multi-parameter task ⇒ may benefit from ML-based methods